Gravitational shockwave collisions in Anti-de Sitter spaces

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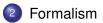
December 21 2015, VIII Black Holes Workshop, IST, Lisboa, Portugal

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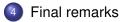
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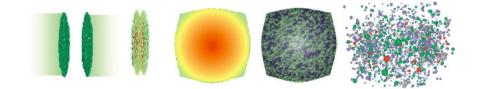






Ultra-relativistic heavy-ion collisions

two nuclei approach, collide, form a QGP, the QGP expands and hadronizes, finally hadrons rescatter and freeze out



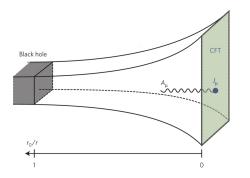
Can we describe all stages in a single framework?

Holography says: Yes! (up to last one)

AdS/CFT

$\mathcal{N}=4$ super-Yang-Mills is dual to IIB string theory on $\textit{AdS}_5 \times \textit{S}_5$

[Maldacena, Witten '98]



 We can learn about strongly coupled phenomena through gravity computations

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AdS/CFT

- $\mathcal{N}=4\;\text{SYM}$
 - conformally invariant
 - no confinement
 - supersymmetric

QCD

- non-conformal
- confinement
- not supersymmetric

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Einstein-Scalar

$$egin{aligned} & R_{\mu
u} - rac{R}{2} g_{\mu
u} = 8\pi T_{\mu
u}\,, \ & \Box \phi = rac{\partial V}{\partial \phi}\,, \end{aligned}$$

where

$$egin{aligned} &8\pi \, T_{\mu
u} = 2 \partial_\mu \phi \partial_
u \phi - g_{\mu
u} \left(g^{lphaeta} \partial_lpha \phi \partial_eta \phi + 2 V(\phi)
ight) \,, \ &V(\phi) = -3 - rac{3}{2} \phi^2 - rac{1}{3} \phi^4 + \left(rac{1}{2 \phi_{
m M}^4} + rac{1}{3 \phi_{
m M}^2}
ight) \phi^6 - rac{1}{12 \phi_{
m M}^4} \phi^8 \,, \end{aligned}$$

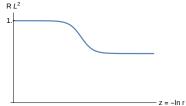
 $\phi_{\rm M}$ is a free parameter

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Scalar field

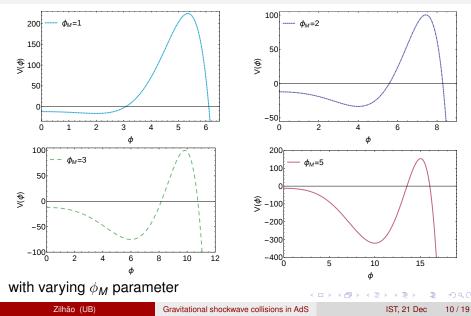
Deforming $\mathcal{N}=$ 4 Super Yang-Mills with a dimension 3 operator $\mathcal O$ dual to the scalar field ϕ

We choose *V* to interpolate between two AdS spaces:



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Potential shapes



Characteristic formulation

D = 5 metric in Eddington-Finkelstein coordinates

$$ds^2 = -Adt^2 + \Sigma^2 \left(e^B dx_\perp^2 + e^{-2B} dz^2\right) + 2dt(dr + Fdz),$$

Schematic evolution equations:

$$\partial_r S = H_S(S, B)$$

 $\partial_t \partial_r B = H_B(B, S, \partial_t B)$

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Advantages of characteristic evolution

- Initial data is free (no elliptic constraints on the data);
- No second time derivatives (therefore smaller number of basic variables);
- Equations have convenient hierarchical structure in which variables are integrated in turn in terms of characteristic data from prior members of the hierarchy.

Initial data

$$ds^{2} = dr^{2} + f(r)h(x_{\perp})dx_{\perp}^{2} - e^{2A(r)}dx_{\perp}dx_{-} + e^{2A(r)}dx_{\perp}^{2}$$

Using the Gaussian profile (*h* height, ω width):

$$h(x_{\perp},z)=he^{-z^2/(2\omega^2)}$$

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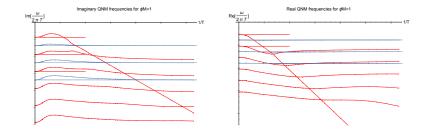






Results

QNMs

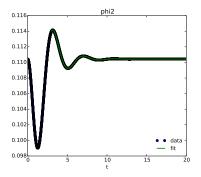


QNM frequencies for $\phi_M = 1$, corresponding to the metric tower and the scalar tower

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Results

QNMs

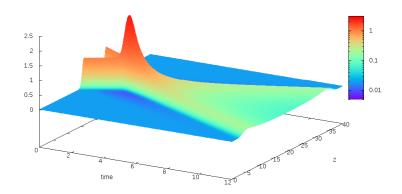


 Results from the code in excellent agreement with perturbative computation

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Energy density



Gravitational shockwave collisions in AdS

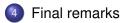
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Zilhão (UB)

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Final Remarks

- Shockwave collisions in AdS spaces provide convenient framework to study heavy-ion collisions
- First simulation of a holographic non-conformal model for heavy ion collisions
- TODO:
 - hydrodynamization time
 - explore parameter space
 - asymmetrical collisions
 - different potentials

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